



Geometric characterization of pointwise slant curves

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Abstract

In the present paper we study the characteristics of pointwise slant curves in a normal almost contact semi-Riemannian three-manifold N^3 . These curves are characterized by the pseudo-Riemannian scalar product between the normal vector at the curve and the reeb vector field of manifold N^3 . In this class of manifolds, curvature and torsion of such curves are determined. The Lancret of slant curves in manifold N^3 is obtained. Additionally, pointwise slant curves with proper mean curvature are characterized.

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1. Introduction

The study of slant curves in contact three-manifolds was started by the authors in [1]. According to [2], slant curves are the generalization of Legendre curves. More precisely, let $(N^3; \varphi, \xi, \eta, g)$ be an almost contact Riemannian three-manifold. Then a smooth unit speed curve $v : J \rightarrow N^3$ is called *slant curve* if $g(\xi, v'(s)) = \cos\vartheta(s) = \text{constant}$, where J is an open interval and $\vartheta : I \rightarrow [0, 2\pi)$ is called

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structural angle (or contact angle). In light of Theorem 3.1 of [1], we obtain for a non-geodesic slant curve v in Sasakian three-manifolds that

$$\text{Lancret}_\pm(v) = \frac{\pm\tau - 1}{\kappa}. \tag{1.1}$$

The equation (1.1) signifies that the curve v is a Legendre helix if and only if the absolute value of its torsion is equal to 1, that is, $|\tau| = 1$. Several authors exhaustively studied and analyzed the geometry of slant curves (see [3–7]). Recently, in [8], the present authors defined pointwise slant curves (abbreviated as PS curves) as a natural generalization of slant curves. In this paper, we investigate how these curves are characterized in N^3 , where N^3 is a normal almost contact semi-Riemannian three-manifold (abbreviated as a. c. s. three-manifold).

The organizational structure of the paper is as follows: the basics of almost contact semi-Riemannian manifolds are given in Sect. 2 and Sect. 3. The characterizations of PS curves in N^3 are obtained in Sect. 4. The curvature and torsion of PS curves which are not geodesic in N^3 are determined in Sect. 5. We derive the necessary and sufficient condition for the PS curve (which is not geodesic) having proper mean curvature vector \mathbb{H} . Examples are also constructed for illustration.

2. Preliminaries

Let the manifold N^{2n+1} of dimension $(2n + 1)$ be C^∞ and paracompact. Let the Lie algebra of vector fields on N^{2n+1} is denoted by $\Xi(N^{2n+1})$ and $\Gamma(F)$ denotes $\Xi(N^{2n+1})$ -module of sections of vector bundle F over the manifold.

The manifold N^{2n+1} is referred to as an *almost contact manifold* if the structure group $\text{GL}_{2n+1}\mathbb{R}$ of TN^{2n+1} (tangent bundle) is reducible to $U(n) \times \{1\}$. Equivalently, if there exists (φ, ξ, η) -structure satisfying

$$\varphi^2 + I = \eta \otimes \xi \quad \text{and} \quad \eta(\xi) = 1, \tag{2.1}$$

where the vector field ξ is called *characteristic* or *Reeb vector field*, φ is an endomorphism, I denotes the identity, and η is a 1-form such that $\eta \wedge (d\eta)^n \neq 0$ known as *contact form*; d is one of the mathematical operators called the exterior differential operator. It is simple to deduce from equation (2.1) that $\eta \circ \varphi = \varphi\xi = 0$ and $\text{rank}(\varphi) = 2n$ [9].

A semi-Riemannian metric g is called compatible with the (φ, ξ, η) -structure if

$$g(\varphi\cdot, \varphi\cdot) + \varepsilon\eta(\cdot)\eta(\cdot) = g(\cdot, \cdot),$$

where g has the signature $(2q + 1, 2n - 2q)$ or $(2q, 2n - 2q + 1)$ depending on whether ξ is spacelike or timelike, respectively, and $\varepsilon^2 = 1$. $(N^{2n+1}; \varphi, \xi, \eta, g)$ is known as an *almost contact semi-Riemannian $(2n + 1)$ -manifold* (abbreviated as a. c. s. $(2n + 1)$ -manifold). Here, $g(\xi, \xi) = \varepsilon$ and $\eta(X) = \varepsilon g(X, \xi)$. This implies that ξ is never lightlike. Let Φ denotes the *fundamental 2-form* then it is given by $\Phi(\cdot, \cdot) = \varepsilon g(\cdot, \varphi\cdot)$. Let the manifold N^{2n+1} further satisfies $d\eta = \Phi$, then it is called a *contact semi-Riemannian manifold*. Let \mathfrak{h} denotes the tensor field defined by $\mathfrak{h} = (1/2)\mathfrak{L}_\xi\varphi$. Then this tensor field plays a crucial role in N^{2n+1} , where \mathfrak{L} denotes the operator of Lie-derivative. Here \mathfrak{h} is self-adjoint and satisfies

$$\nabla_{\xi\varphi} = 0, \quad \nabla\xi = -\varphi \circ \mathfrak{h} - \varepsilon\varphi \mathfrak{h}(\xi) = \text{trace}(\mathfrak{h}) = 0, \quad \varphi \circ \mathfrak{h} = -\mathfrak{h} \circ \varphi,$$

where ∇ is a Levi-Civita connection. For more details about the geometry of the contact semi-Riemannian manifold, we refer to [10–12].

Let us consider the product manifold $N^{2n+1} \times \mathbb{R}$: $\left(Z, \zeta \frac{d}{dt} \right)$ is an arbitrary tangent vector, ζ

is a smooth function on $N^{2n+1} \times \mathbb{R}$, t is the standard coordinate on \mathbb{R} , and $Z \in \Gamma(TN^{2n+1})$.

The almost complex structure J on this direct product is given as follows:

$$J\left(Z, \xi \frac{d}{dt}\right) = \left(\varphi Z - \xi \eta(Z) \frac{d}{dt}\right).$$

Then N^{2n+1} is called *normal* if and only if

$$d\eta(\cdot, \cdot)\xi + \frac{1}{2}[\varphi, \varphi](\cdot, \cdot) = 0,$$

where $[\varphi, \varphi]$ denotes the *Nijenhuis torsion*, and it is given as follows:

$$[\varphi, \varphi](\cdot, \cdot) = [\varphi \cdot, \varphi \cdot] - \varphi[\varphi \cdot, \cdot] + \varphi^2[\cdot, \cdot] - \varphi[\cdot, \varphi \cdot]$$

(see [9, 11]).

3. Normal a. c. s. Three-Manifolds

In the present paper, we restrict ourselves to dimension three. Analogous to [13], we give some results related to this case. If we consider N^3 to be an a. c. s. three-manifold, we find that

$$(\nabla_{Z_1} \varphi)Z_2 = -\eta(Z_2)\varphi\nabla Z_1\xi + \varepsilon g(\varphi\nabla Z_1\xi, Z_2)\xi, \tag{3.1}$$

where $Z_1, Z_2 \in \Gamma(TN^3)$.

Proposition 3.1. *In an a. c. s. three-manifold N^3 , the following conditions are mutually equivalent:*

- (i) manifold M^3 is normal;
- (ii) $\varphi\nabla Z\xi = \nabla\varphi Z\xi$;
- (iii) $\nabla Z\xi = \varepsilon\beta(Z - \eta(Z)\xi) - \varepsilon\alpha\varphi Z$.

Here $Z \in \Gamma(TN^3)$, α and β being smooth functions on N^3 for which we have

$$2\alpha = \text{trace}\{Z \rightarrow \varphi\nabla Z\xi\}, \quad 2\beta = \text{trace}\{Z \rightarrow \nabla Z\xi\}.$$

From equation (3.1) and Proposition 3.1, we find that

$$(\nabla_{Z_1} \varphi)Z_2 = \beta(g(\varphi Z_1, Z_2)\xi - \varepsilon\eta(Z_2)\varphi Z_1) + \alpha(g(Z_1, Z_2)\xi - \varepsilon\eta(Z_2)Z_1). \tag{3.2}$$

Moreover, manifold N^3 satisfies

$$\xi(\alpha) + 2\varepsilon\alpha\beta = 0.$$

Therefore, $\beta = 0$ if α is a non-zero constant. Analogous to [5], N^3 is called

- *Cosymplectic semi-Riemannian manifold* if $\alpha = \beta = 0$;
- *quasi-Sasakian semi-Riemannian manifold* if $\beta = 0$ and $\xi(\alpha) = 0$;
- *β -Kenmotsu semi-Riemannian manifold* if $\alpha = 0$ and β is a non-zero constant.

In addition, N^3 is said to be a *Sasakian semi-Riemannian manifold* if $\alpha = 1, \beta = 0$ and *Kenmotsu semi-Riemannian manifold* if $\alpha = 0, \beta = 1$. Now, we give examples of normal a. c. s. three-manifolds.

Example 3.2. Consider the standard Cartesian coordinates on \mathbb{R}^3 as (x, y, z) , 1-form η is given by $\eta = ydx + dz$, $\xi = \partial z$ and the endomorphism φ is defined by $\varphi\partial_x = \partial_y, \varphi\partial_y = y\partial_z - \partial_x, \varphi\partial_z = 0$, where $\partial_x = \frac{\partial}{\partial x}, \partial_y = \frac{\partial}{\partial y}$ and $\partial_z = \frac{\partial}{\partial z}$. Then $\varphi^2 + \mathcal{I} = \eta \otimes \xi$ and $\eta(\xi) = 1$ are obtained. Therefore, the (φ, ξ, η) -structure is almost contact. Further, by simple computations, we find that the (φ, ξ, η) -structure is normal.

Let $\mathcal{N}_\varepsilon^3 := \mathbb{R}_1^2 \times \mathbb{R}_+ \subset \mathbb{R}_1^3$ and normal a. c. s. structure (φ, ξ, η, g) is restricted to $\mathcal{N}_\varepsilon^3$, where $g = \varepsilon\eta \otimes \eta + z^2(dx^2 + dy^2)$. Then, we have $g(\partial_x, \partial_x) = \varepsilon y^2 + z^2, g(\partial_y, \partial_y) = z^2, g(\partial_z, \partial_z) = \varepsilon, g(\partial_x, \partial_y) = g(\partial_x, \partial_z) = 0$,

$g(\partial_x, \partial_z) = g(\partial_z, \partial_x) = \varepsilon y$, $g(\partial_y, \partial_z) = g(\partial_z, \partial_y) = 0$, where $\varepsilon^2 = 1$. Using φ and g , we have $g(\varphi Z_1, \varphi Z_2) = g(Z_1, Z_2) - \varepsilon \eta(Z_1)\eta(Z_2)$ and $\eta(Z_1) = \varepsilon g(Z_1, \xi)$, and thus $(N_\varepsilon^3; \varphi, \xi, \eta, g)$ is a normal a. c. s. three-manifold. For ∇ with respect to g , we have

$$\begin{aligned} \nabla_{\partial_x} \partial_x &= \frac{y}{z} \partial_x - \frac{\varepsilon y}{z^2} \partial_y - \left(\frac{z^2 + \varepsilon y^2}{\varepsilon z} \right) \partial_z, \nabla_{\partial_x} \partial_y = \nabla_{\partial_y} \partial_x = \frac{\varepsilon y}{2z^2} \partial_x + \left(\frac{z^2 - \varepsilon y^2}{2z^2} \right) \partial_z, \\ \nabla_{\partial_x} \partial_z &= \nabla_{\partial_z} \partial_x = \frac{1}{z} \partial_x - \frac{\varepsilon}{2z^2} \partial_y - \frac{y}{z} \partial_z, \nabla_{\partial_y} \partial_y = \frac{y}{z} \partial_x - \left(\frac{z^2 + \varepsilon y^2}{\varepsilon z^2} \right) \partial_z, \\ \nabla_{\partial_y} \partial_z &= \nabla_{\partial_z} \partial_y = \frac{\varepsilon}{2z^2} \partial_1 + \frac{1}{z} \partial_y - \frac{\varepsilon y}{2z^2} \partial_z, \nabla_{\partial_z} \partial_z = 0. \end{aligned}$$

Using the expressions above and the equation (3.2), we have $\beta = \frac{\varepsilon}{z}$ and $\alpha = \frac{1}{2z^2}$.

Example 3.3. Let $\mathcal{K}_\varepsilon^3 := \mathbb{R}_1^3$ with (x, y, z) as standard Cartesian coordinates, $\eta = dz$, $\xi = \partial_z$, endomorphism φ satisfies: $\varphi \partial_x = \partial_y$, $\varphi \partial_y = -\partial_x$, $\varphi \partial_z = 0$ and metric tensor is given as $g = \varepsilon \eta \otimes \eta + \exp(2z)(dx^2 + dy^2)$, where $\partial_x = \frac{\partial}{\partial x}$, $\partial_y = \frac{\partial}{\partial y}$ and $\partial_z = \frac{\partial}{\partial z}$. Then, by making straightforward computations, we find that $(\mathcal{K}_\varepsilon^3; \varphi, \xi, \eta, g)$ is a normal a. c. s. three-manifold. For ∇ with respect to this g , we have

$$\begin{aligned} \nabla_{\partial_x} \partial_x &= \nabla_{\partial_y} \partial_y = -\varepsilon \exp(2z) \partial_z, \nabla_{\partial_x} \partial_z = \nabla_{\partial_z} \partial_x = \partial_x, \\ \nabla_{\partial_y} \partial_z &= \nabla_{\partial_z} \partial_y = \partial_y, \nabla_{\partial_x} \partial_y = \nabla_{\partial_y} \partial_x = \nabla_{\partial_z} \partial_z = 0. \end{aligned} \tag{3.3}$$

By the virtue of partial differential equations (3.3) and equation (3.2), we obtain $\beta = \varepsilon$ and $\alpha = 0$. Therefore, $\mathcal{K}_\varepsilon^3$ is a 3-dimensional ε -Kenmotsu manifold. Furthermore, $\mathcal{K}_\varepsilon^3$ is a warped product $\mathbb{R} \times_f \mathbb{F}^2$ where warping function f is given by $f(z) = \exp(\varepsilon z)$.

For more information on warped geometry, we may refer to [14].

4. Pointwise Slant Curves

Let N^3 be a normal a. c. s. three-manifold with Levi-Civita connection ∇ , $v : I \rightarrow N^3$ be an unit speed curve in N^3 , I being an open interval. Then v is called a *Frenet curve* if the Frenet frame $\{T := v', N, B\}$ of v satisfies (*Frenet–Serret formulas*) [7, p. 968]:

$$\nabla_{T'} T = \kappa N, \nabla_{T'} N = \varepsilon \tau B - \kappa T \text{ and } \nabla_{T'} B = -\tau N, \tag{4.1}$$

where $\kappa = |\nabla_{T'} T|$ and τ are denote the curvature and torsion of v , respectively. The vectors T, B and N are known as the tangent, binormal and principal normal of v , respectively. The curve v is called geodesic if $\nabla_{v'} v' = 0$ and it is not geodesic if $\kappa > 0$ everywhere on I .

Following [8], we give

Definition 4.1. Let N^3 be a normal a. c. s. three-manifold and $v : J \rightarrow N^3$ be a Frenet curve. Let $\rho : I \rightarrow I_1 \subseteq \mathbb{R}$ be a smooth function, where $I_1 = [-1, 1]$ or $I_1 = [0, \infty)$. Then v is said to be a *pointwise slant curve* (abbreviated as *PS curve*) if $\eta(v') = \rho$. We call ρ , a *slant function*. In particular, v is *slant curve* if $\rho = \text{constant}$ [4] and if $\rho = 0$ it is *Legendre curve* ([2, 9]). The PS curve is said to be *proper*, if neither $\rho = 0$ nor $\rho = \text{constant}$.

Remark 4.2. For a PS curve v in N^3 , we have

$$\rho = \varepsilon g(\xi, v'(s)). \tag{4.2}$$

If the characteristic vector field is timelike, then $\rho = \sinh \vartheta_1(s)$, where angle $\vartheta_1 : I \rightarrow [0, \infty)$ is called *Lorentzian timelike* between v' and characteristic vector field [15]. In this case, $\rho \in [0, \infty)$. Further, let

$\{v', \xi\}$ span a spacelike vector subspace, and if characteristic vector field is spacelike, then $\rho = \cos\vartheta_2(s)$, where $\vartheta_2 : I \rightarrow [0, 2\pi)$ is the contact angle of v [4]. In this case, $\rho \in [-1, 1]$.

Using Definition 4.1, we find that a curve $v(s) = (v_1(s), v_2(s), v_3(s))$ in $\mathcal{N}_\varepsilon^3$ is a PS if and only if

$$\begin{cases} v'_1 v_2 + v'_3 = \rho, \\ v_3^2 (v_1'^2 + v_2'^2) = 1 - \varepsilon\rho^2, \end{cases}$$

where ρ is a smooth function. It can be easily seen that

$$v' = v'_1 \partial_1 + v'_2 \partial_2 + (\rho - v'_1 v^2) \partial_3 \text{ and } \varphi v' = -v'_2 \partial_1 + v'_1 \partial_2 + v_2 v'_2 \partial_3.$$

Here, it is important to mention that every unit speed curve in N^3 is not necessarily a PS curve, for instance, consider the following curve in \mathcal{N}_{-1}^3 :

$$\gamma(s) = \left(-\sqrt{5}s, \frac{2}{\sqrt{5}}, 1\right), s \in \mathbb{R}.$$

Then we have $g(\gamma', \gamma) = 1$. Here $\rho = -2$, this implies that $\sinh \vartheta_1(s) = -2$, which is not possible value for the above defined smooth function ρ .

After taking covariant differentiation of equation (4.2) along v , we get

$$\rho' = \varepsilon g(\xi, \kappa N) + g(v', -\alpha\varphi v' + \beta(v' - \rho\xi)) = \beta(1 - \varepsilon\rho^2) + \kappa\eta(N). \tag{4.3}$$

The interpretation of ξ in terms of the Frenet frame of v provides

$$\varepsilon\eta(N)^2 + \eta(B)^2 = 1 - \varepsilon\rho^2. \tag{4.4}$$

Using equations (4.3) and (4.4), we have the following characterization result for the PS curve:

Proposition 4.3. *Let $v : J \rightarrow N^3$ be a non-geodesic curve. Then v is a PS curve if and only if*

$$\eta(N) = \frac{\rho' - \beta(1 - \varepsilon\rho^2)}{\kappa}. \tag{4.5}$$

Therefore, a necessary condition for v to be a PS curve is

$$\varepsilon\eta(N)^2 \leq (1 - \varepsilon\rho^2), \tag{4.6}$$

only if $\rho \neq \pm 1$.

Remark 4.4.

- (i) The characterization (4.5) (as well as (4.6)) is independent of the Sasakian part, *i.e.*, does not depend on α . Thus for a (semi-Riemannian) quasi-Sasakian manifold, this expression gives $\eta(N) = \frac{\rho'}{\kappa}$. Particularly, for ρ -slant curve v which is not geodesic in quasi-Sasakian 3-manifold, we have $\eta(N) = 0$ [1].
- (ii) From (4.6), equality yields $\eta(N)^2 = 1 - \rho^2$ (then $\eta(T) = \rho$ and $\eta(B) = 0$) only if ξ is spacelike. Particularly, v is a Legendre curve with $\kappa = \beta|v$ and $N = -\xi$.
- (iii) For a slant curve v with $\varepsilon = 1$, equation (4.5) provides $\eta(N) = -\frac{\beta}{\kappa} \sin^2 \vartheta_1$ [4].

Let v be a PS curve in N^3 . Consider $v', \varphi v', \xi$ such that $g(\xi, \xi) = \varepsilon$, $g(v', v') = 1$, $g(v', \xi) = \varepsilon\rho$, $g(\varphi v', \varphi v') = 1 - \varepsilon\rho^2$ and $g(v', \varphi v') = g(\xi, \varphi v') = 0$; ρ being a smooth function. Then, the set $\{v', \varphi v', \xi\}$ is linearly independent, forms a basis of $\mathcal{T}_{v(s)}N^3$ for every $s \in I$ if and only if $m = \sqrt{|1 - \varepsilon\rho^2|} \neq 0$. Now, we can define orthonormal vector fields as:

$$B_1 = v', \quad B_2 = \frac{\varphi v'}{m}, \quad B_3 = \frac{\xi - \varepsilon\rho v'}{m}, \tag{4.7}$$

where $m = \sqrt{|1 - \varepsilon\rho^2|}$, $g(B_3, B_3) = \varepsilon$ and $g(B_1, B_1) = g(B_2, B_2) = 1$.

Here, $\{v', \varphi v', \xi\}$ is linearly dependent if and only if $v' = \varepsilon\varphi v' + \rho\xi$ or $v' = \rho\xi$. Furthermore, if $\{v', \varphi v', \xi\}$ is linearly dependent, then $|\rho| = 1$ and ξ is spacelike. This implies that v is necessarily a geodesic. Therefore, we must have $m \neq 0$ for the non-geodesic curve v . The decomposition of ξ with respect to $\{B_1, B_2, B_3\}$ is as follows:

$$\xi = \varepsilon(mB_3 + \rho B_1). \tag{4.8}$$

Remark 4.5. We define the *Lancret coefficient* of a PS curve v in N^3 which is not geodesic by

$$\text{Lancret}(v) = \frac{\rho}{m}. \tag{4.9}$$

The insight for above definition is that for $\varepsilon = 1$, the above expression yields $\text{Lancret}(v) = \frac{\cos \vartheta}{|\sin \vartheta|}$, where $\vartheta = \text{constant}$, analogous to contact geometry [3].

5. Main Results

Let $\nabla_v B_1 = a_1 B_1 + b_1 B_2 + c_1 B_3$, where a_1, b_1 and c_1 are any C^∞ functions. Then $a_1 = g(\nabla_v B_1, B_1) = 0$, $b_1 = g(\nabla_v B_1, B_2) = \delta m$ and $-\varepsilon c_1 = g(B_1, \nabla_v B_3)$, where $\delta = \frac{1}{m^2} g(\nabla_v v', \varphi v')$. Using equation (3.2), we get $g(B_1, \nabla_v B_3) = \varepsilon \left(\beta m - \frac{\rho'}{m} \right)$. Thus, $\nabla_v B_1 = \delta m B_2 - \left(\beta m - \frac{\rho'}{m} \right) B_3$. Analogy to this, we can find $\nabla_v B_2$ and $\nabla_v B_3$. This provides the following result:

Lemma 5.1. *Let $v : J \rightarrow N^3$ be a PS curve which is not geodesic. Then, we have*

$$\nabla_v B_1 = m\delta B_2 + \left(\frac{\rho'}{m} - m\beta \right) B_3, \tag{5.1}$$

$$\nabla_v B_2 = -m\delta B_1 + (\alpha + \delta\rho) B_3, \tag{5.2}$$

$$\nabla_v B_3 = \varepsilon \left(\left(m\beta - \frac{\rho'}{m} \right) B_1 - (\delta\rho + \alpha) B_2 \right), \tag{5.3}$$

where

$$m = \sqrt{|1 - \varepsilon\rho^2|} \text{ and } \delta = \frac{1}{m^2} g(\nabla_v v', \varphi v'). \tag{5.4}$$

Theorem 5.2. *Let $v : J \rightarrow N^3$ be a PS curve which is not geodesic. Then expression for curvature κ and torsion τ of v are as follow:*

$$\left\{ \begin{array}{l} \kappa = m \sqrt{\varepsilon \left(\beta - \frac{\rho'}{m^2} \right)^2 + \delta^2}, \\ \tau = \pm \left(\alpha + \rho\delta + \frac{(\beta\delta' - \beta'\delta) - 2 \left(\frac{\rho'\delta'}{m^2} \right) + \left(\frac{\delta\rho'}{m^2} \right)'}{\varepsilon \left(\beta - \frac{\rho'}{m^2} \right)^2 + \delta^2} \right), \end{array} \right. \tag{5.5}$$

where m and δ are given in equation (5.4).

Proof. From equation (5.1) and by computation of length of $\nabla_{v'}v'$ i.e. $\|\nabla_{v'}v'\|$, we receive κ . In light of equations (4.1) and (5.1), we get

$$N = \frac{m\delta}{\kappa} B_2 - \left(\frac{m\beta}{\kappa} - \frac{\rho'}{m\kappa} \right) B_3.$$

Let $a = m\delta$ and $b = \left(m\beta - \frac{\rho'}{m} \right)$. Then

$$\begin{aligned} \nabla_{v'}N &= \left(-\frac{a^2}{\kappa} - \frac{\varepsilon b^2}{\kappa} \right) B_1 + \left(\frac{a'\kappa - a\kappa'}{\kappa^2} + \frac{\varepsilon b}{\kappa} \left(\alpha + \frac{a\rho}{m} \right) \right) B_2 + \left(-\frac{\kappa b' - \kappa'b}{\kappa^2} + \frac{a}{\kappa} \left(\frac{a\rho}{m} + \alpha \right) \right) B_3 \\ &= -\kappa B_1 + \left(\frac{\varepsilon b B_2 + a B_3}{\kappa} \right) \left(\alpha + \frac{\rho a}{m} + \frac{b a' - b' a}{\kappa^2} \right). \end{aligned} \tag{5.6}$$

Using Frenet–Serret formulas, we obtain from equation (5.6) that $\tau = \pm \left(\frac{\rho a}{m} + \frac{b a' - b' a}{\kappa^2} + \alpha \right)$. Hence, it completes the proof.

Now, we can give the following result as a corollary of the above theorem:

Corollary 5.3. *Let v be a slant curve in N^3 which is not geodesic. Then the expressions for curvature and torsion of v are as follows:*

$$\begin{cases} \kappa = m\sqrt{|\varepsilon\beta^2 + \delta^2|}, \\ \tau = \pm \left(\rho\delta + \alpha + \frac{\beta\delta' - \beta'\delta}{\delta^2 + \varepsilon\beta^2} \right). \end{cases} \tag{5.7}$$

Then the associated Lancret of v for $\delta \neq 0$, is given by

$$\text{Lancret}_{\pm}(v) = \frac{(\beta'\delta - \beta\delta')|\delta^2 + \varepsilon\beta^2|^{\frac{1}{2}} - (\alpha \mp \tau)|\delta^2 + \varepsilon\beta^2|^{\frac{1}{2}}}{\kappa\delta}, \tag{5.8}$$

where m and δ are given in equation (5.4).

Proof. Since $\rho' = 0$, i.e., v is slant. Thus, equation (5.5) leads to (5.7) and equation(5.8) follows directly from (4.9) and (5.7). Hence, it completes the proof.

Consider $\xi = p_1T + p_2N + p_3B$, where p_1, p_2 , and p_3 are any smooth functions. By the virtue of equations (4.1), (4.7), (5.1)-(5.3), and (5.6), we can readily compute p_1, p_2 , and p_3 . This provides the following result:

Proposition 5.4. *Let $v : J \rightarrow N^3$ be a PS curve which is not geodesic. Then, the decomposition of ξ is expressed as follows*

$$\xi = \varepsilon(\rho\kappa T + (\rho' - m^2\beta)N + m^2\text{sgn}(\tau)\delta B)/\kappa,$$

where $\{T, N, B\}$ denotes the Frenet frame of v and m, δ are given in equation (5.4).

Now, we present the κ and τ values of a non-geodesic PS curve in some subclasses of N^3 .

Corollary 5.5. *Let $v : J \rightarrow N^3$ be a PS curve which is not geodesic.*

(a) *Let N^3 be a Cosymplectic semi-Riemannian manifold then we have*

$$\begin{cases} \kappa = m \sqrt{\left| \varepsilon \left(\frac{\rho'}{m^2} \right)^2 + \delta^2 \right|}, \\ \tau = \pm \left(\rho\delta - \frac{2 \left(\frac{\rho'\delta'}{m^2} \right) - \left(\frac{\rho'\delta}{m^2} \right)'}{\varepsilon \left(\frac{\rho'}{m^2} \right)^2 + \delta^2} \right). \end{cases} \tag{5.9}$$

(b) *Let N^3 be a quasi-Sasakian semi-Riemannian manifold then we have*

$$\begin{cases} \kappa = m \sqrt{\left| \varepsilon \left(\frac{\rho'}{m^2} \right)^2 + \delta^2 \right|}, \\ \tau = \pm \left(\alpha + \rho\delta - \frac{2 \left(\frac{\rho'\delta'}{m^2} \right) - \left(\frac{\rho'\delta}{m^2} \right)'}{\varepsilon \left(\frac{\rho'}{m^2} \right)^2 + \delta^2} \right). \end{cases}$$

(c) *Let N^3 be a β -Kenmotsu semi-Riemannian manifold then we have*

$$\begin{cases} \kappa = m \sqrt{\left| \varepsilon \left(\beta - \frac{\rho'}{m^2} \right)^2 + \delta^2 \right|}, \\ \tau = \pm \left(\rho\delta + \frac{\beta\delta' - 2 \left(\frac{\rho'\delta'}{m^2} \right) + \left(\frac{\rho'\delta}{m^2} \right)'}{\varepsilon \left(\beta - \frac{\rho'}{m^2} \right)^2 + \delta^2} \right), \end{cases}$$

where m and δ are given in equation (5.4).

Below, we give certain proper PS curves in $\mathcal{N}_\varepsilon^3$:

Example 5.6. Let in \mathcal{N}_1^3

$$v^1(s) = (s, 0, \sin s), \quad s \in (0, 2\pi).$$

Then the curve v^1 is a proper PS curve in \mathcal{N}_1^3 . Here, we have $\alpha(v^1(s)) = \frac{1}{2} \csc^2 s$, $\beta(v^1(s)) = \csc s$, $\delta(s) = -\cot s \csc s$, $\rho = \cos s$,

$$\kappa = \left(1 + 3 \sin^2 s \right)^{\frac{1}{2}} \csc s \text{ and } \tau = \left| \frac{\cot^2 s}{2\kappa^2} - 1 \right|.$$

The Euclidean image of v^1 is depicted in Figure 1. Moreover, some of the particular cases of v^1 are striking, these are portrayed in Figures 2, 3, and 4.

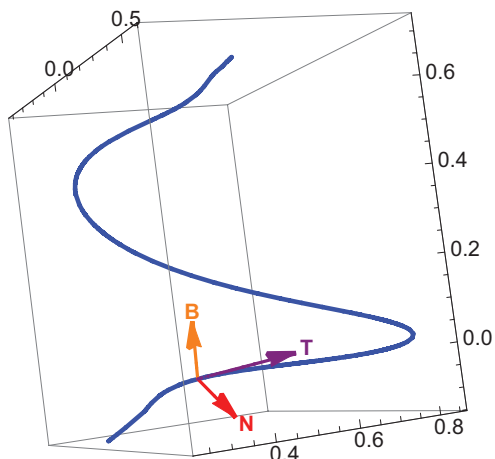


Figure 1: $\rho = \cos s$

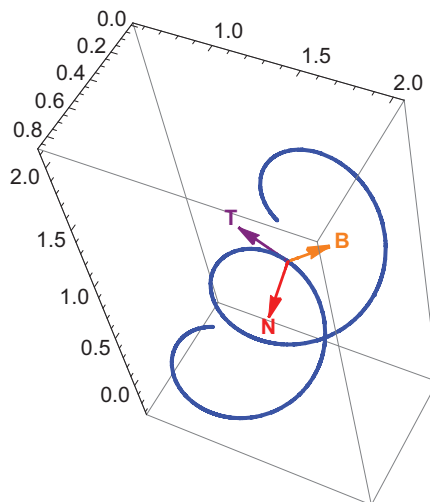


Figure 2: $\rho = \frac{\sqrt{3}}{2}$.

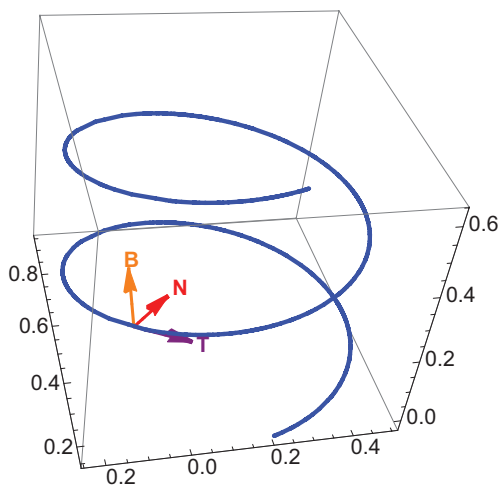


Figure 3: $\rho = \frac{1}{2}$.

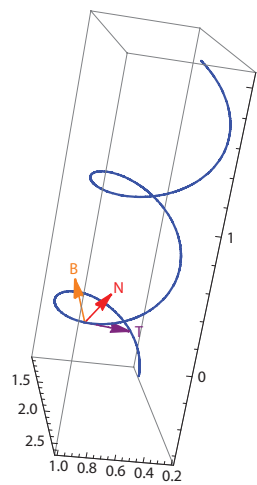


Figure 4: $\rho = 0$.

Example 5.7. Let in \mathcal{N}_{-1}^3

$$v^2(s) = (k, s, \cosh s), \quad s \in (0, \infty), k \in \mathbb{R}.$$

Then v^2 is a proper PS curve in \mathcal{N}_{-1}^3 . Here, we find $\beta(v_2(s)) = -\operatorname{sech} s$, $\delta(s) = \operatorname{sech} s(-s + \tanh s)$, $\alpha(v^2(s)) = \frac{1}{2} \operatorname{sech}^2 s$, $\rho = \sinh s$,

$$\kappa = \sqrt{|4 - (s - \tanh s)^2|} \text{ and } \tau = \left| (\tanh s)(s - \tanh s) - \frac{1}{2} \operatorname{sech}^2 s + \frac{2 \tanh^2 s}{4 - (s - \tanh s)^2} \right|.$$

From Definition 4.1, the necessary and sufficient conditions for $v(s) = (v_1(s), v_2(s), v_3(s))$ to be a PS curve in $\mathcal{K}_\varepsilon^3$ are given by

$$\begin{cases} v_3' = \rho, \\ v_1'^2 + v_2'^2 = \exp(-2v_3)m^2. \end{cases} \tag{5.10}$$

Now, we have

$$v' = v'_1 \partial_x + v'_2 \partial_y + \rho \partial_z \text{ and } \varphi v' = -v'_y \partial_x 1 + v'_1 \partial_y.$$

From (5.10), we find

$$v'_1 = m \exp(-v_3) \cos \zeta(s), v'_2 = m \exp(-v_3) \sin \zeta(s), v'_3 = \rho,$$

where $\zeta \in C^\infty(I)$. This leads to the following result:

Proposition 5.8. *Let $v : J \rightarrow \mathcal{K}_\varepsilon^3$ be a PS curve in $\mathcal{K}_\varepsilon^3$ which is not geodesic. Then v can be expressed as follows:*

$$v(s) = \left(\int_{s_0}^s \Psi(t) \Omega(t) m dt, \int_{s_0}^s \rho(t) dt \right),$$

where $\Psi(s) = (\cos \zeta(s), \sin \zeta(s))$ is parametrization of circle \mathbb{S}^1 , $\Omega(s) = \exp\left(-\int_{s_0}^s \rho(t) dt\right)$ and ζ is a smooth function on I .

By straightforward computations, we get

$$\begin{cases} \kappa &= m \sqrt{\left| \varepsilon \left(1 - \frac{\varepsilon \rho'}{m^2} \right)^2 + \zeta'^2 \right|}, \\ \tau &= \pm \left(\rho \zeta' + \frac{\varepsilon \zeta'' - 2 \left(\frac{\zeta' \rho'}{m^2} \right) + \left(\frac{\rho' \zeta'}{m^2} \right)'}{\zeta'^2 + \varepsilon \left(1 - \frac{\varepsilon \rho'}{m^2} \right)^2} \right). \end{cases}$$

Let us consider a PS curve $v : J \rightarrow N^3$ in N^3 . Then we have

$$\mathbb{H} = \nabla_{v'} v',$$

where \mathbb{H} denotes the mean curvature vector field. Then PS curve v is known as a curve with proper \mathbb{H} , if we have a C^∞ function λ such that

$$\Delta \mathbb{H} = \lambda \mathbb{H}. \tag{5.11}$$

Here Δ denotes the Laplace operator and it is explicitly given by

$$\Delta = -\nabla_{v'} \nabla_{v'}.$$

If $\lambda = 0$ then PS curve v is said to be a curve with harmonic \mathbb{H} ([4, 6]). Using Frenet-Serret formulas, (5.11) can be rewritten as

$$-3\kappa \kappa' v' + \varepsilon(2\kappa' \tau + \kappa \tau') B + (\kappa'' - \kappa^3 - \varepsilon \kappa \tau^2) N = -\kappa \omega N.$$

This implies that the relation $\Delta \mathbb{H} = \lambda \mathbb{H}$ holds if and only if $\kappa \tau' = 0$ and $\varepsilon \tau^2 \kappa + \kappa^3 = \lambda \kappa$. This yields the following remark:

Remark 5.9. Let v be a PS curve in N^3 . Then it holds $\Delta \mathbb{H} = \lambda \mathbb{H}$ if and only if PS curve v is either geodesic or helix satisfying

$$\varepsilon \tau^2 + \kappa^2 = \lambda. \tag{5.12}$$

In view of the above remark, it can be easily seen that the PS curve which is not geodesic with harmonic \mathbb{H} does not exist in N^3 . In light of equations (5.5), (5.11) and (5.12), we provide the following proposition which generalizes result 3.7 of [3] and result 3 of [4].

Proposition 5.10. *If we consider a PS curve v in N^3 which is not geodesic. Then v is having proper \mathbb{H} if and only if it is helix such that $\lambda = \left(\alpha + \rho\delta + \frac{\rho}{\kappa^2}\right)\varepsilon + \kappa^2$, where δ and m are given in (5.4) and $\varpi = (\beta\delta' - \beta'\delta)m^2 - 2\delta'\rho' + m^2\left(\frac{\rho'}{m^2}\right)'$.*

Example 5.11. Let in \mathcal{N}_1^3

$$v^3(s) = \left(1, \ln s, \frac{s}{\sqrt{2}}\right), s \in (0, \infty),$$

then v^3 is a slant Frenet curve in \mathcal{N}_1^3 . Here, we find $\rho = \frac{1}{\sqrt{2}}$ (that is, $\vartheta = \frac{\pi}{4}$), $\beta(v^3(s)) = \frac{\sqrt{2}}{s}$, $\alpha(v^3(s)) = \frac{1}{s^2}$, $\delta(s) = -\frac{\sqrt{2}}{s^2} \ln se \neq 0$,

$$\kappa = \frac{\sqrt{s^2 + (\ln se)^2}}{s^2} \text{ and } \tau = \left| \left(\frac{s^2}{(\ln se)^2 + 1} - 1 \right) \frac{\ln s}{s^2} \right|.$$

Using (5.8), we find Lancret (v^3) = 1. Now, we have that

$$\tau^2 + \kappa^2 = \frac{(2 \ln se - (\ln s)^2)}{s^2(1 + (\ln se)^2)} + \frac{(2 \ln s \ln se + 1)}{s^4} + \frac{(\ln s)^2}{(1 + (\ln se)^2)^2}.$$

From the above equation, we get $\tau^2 + \kappa^2 = \text{non-constant}$. Therefore, by the consequence of Proposition 5.10, we find v^3 is without proper \mathbb{H} .

Example 5.12. Let in \mathcal{N}_{-1}^3

$$v^4(s) = \left(\sqrt{3}s, 1, \frac{2}{\sqrt{3}}\right), s \in \mathbb{R},$$

then v^4 is slant curve in \mathcal{N}_{-1}^3 . For v^4 , we have $\rho = \sqrt{3}$ (that is, $\vartheta = \arcsin h\sqrt{3}$), $\alpha(v^4(s)) = \frac{3}{8}$, $\delta(s) = -\frac{3\sqrt{3}}{4} \neq 0$, $\beta(v^4(s)) = -\frac{\sqrt{3}}{2}$, $\kappa = \frac{\sqrt{15}}{2}$ and $\tau = \frac{21}{8}$. Using equation (5.8), we get Lancret (v^4) = $\frac{\sqrt{3}}{2}$. Now, we obtain that $\kappa^2 - \tau^2 = \text{constant}$. Therefore, in light of Proposition 5.10, we have v^4 is helix and having proper \mathbb{H} with $\lambda = -\frac{201}{64}$.

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